

APPARENT MOTION OF THE ICRF SOURCES

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- *International Celestial Reference Frame*

1. Sources and sky coverage
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3. Kinematical principle

- *Apparent motion of the ICRF sources*

- Observational data
- Angular spectrum representation
- Dipole harmonic, its interpretation
- Apparent motion in the field of a Cosmological Gravitational Wave
- Expected apparent motion and opportunity of observation

What is the ICRF?

A realization of the ICRF consists of a set of precise coordinates of extragalactic radio sources (accuracy is better than $250 \mu\text{as}$ in fact it tends to $50 \mu\text{as}$):

- a) quasars,
- b) BL Lac type objects,
- c) active galactic nuclei (AGN).

The reason is that the proper motion of these radio sources is expected to be negligibly small because of their remoteness.

Expected proper motion is 10 nas/yr

Expected motion and apparent motion

Proper motion of source
(linear motion, constant flux)

$$Z = 1, V = 100 \text{ km/s} \rightarrow \mu \sim 0.01 \text{ } \mu\text{as/y}$$

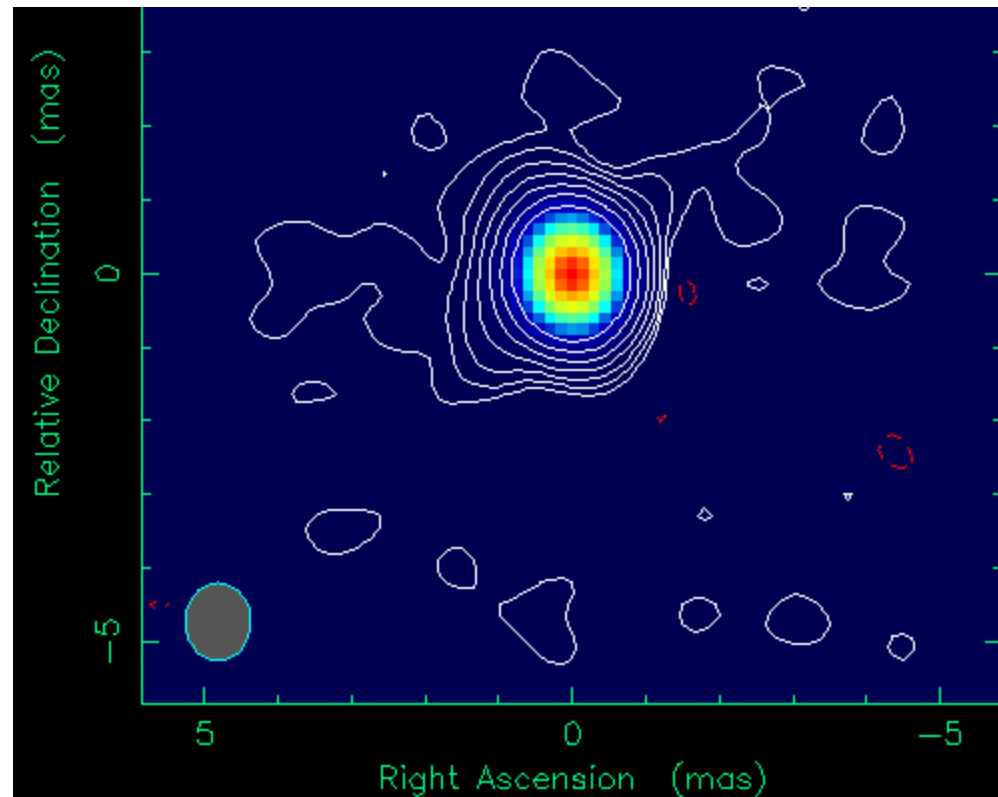
1044+719

$z=1.150$,

$\text{scale} = 8.286 \text{ kpc/''}$

$|\mu| \approx (0.043 \pm 0.002) \text{ mas/yr}$,

$V_T \approx 1.2c$ (apparent)



The first adopted catalogue

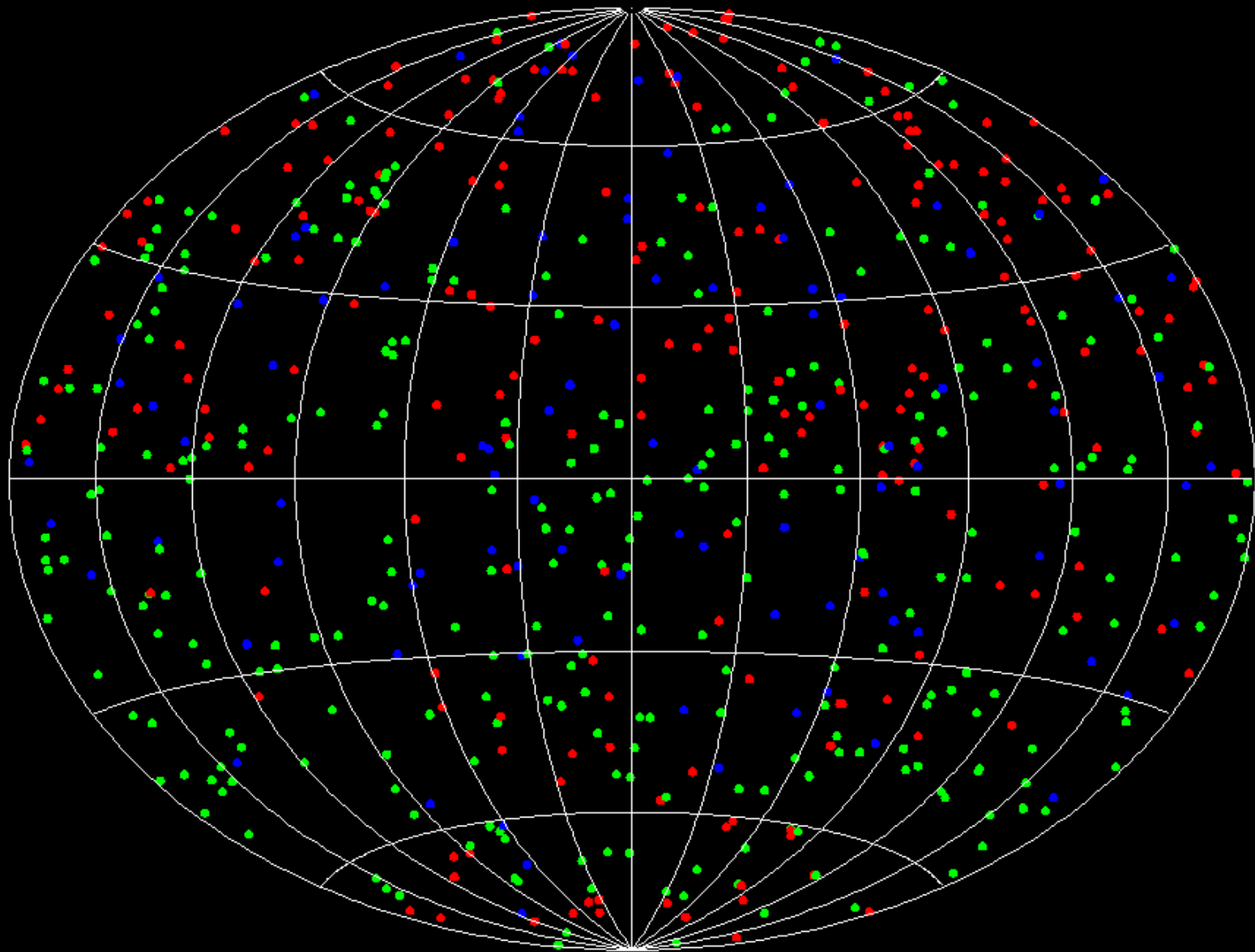
(**608** sources).

- The first realization of the ICRF (catalog) was constructed in 1995 by a reanalysis of the VLBI observations (Ma et al. 1998).
- **212** of these are defining sources providing a core of the ICRF (Ma et al. 1998). The estimated source position uncertainty for the 'defining' sources is about 0.25 mas.
- **294** 'candidate' sources have fewer observations.
- **102** 'other' sources were added to densify the ICRF.

The radio source positions for the current ICRF realization were obtained from the analysis of observations from 1979 to mid 1995.

Two extensions have been announced since 1995 covering observations at 1994 through 2002 (Fey et al. 2001, 2004). Positions for **109** new radio sources were added to the list of the initial ICRF catalogue.

ICRF

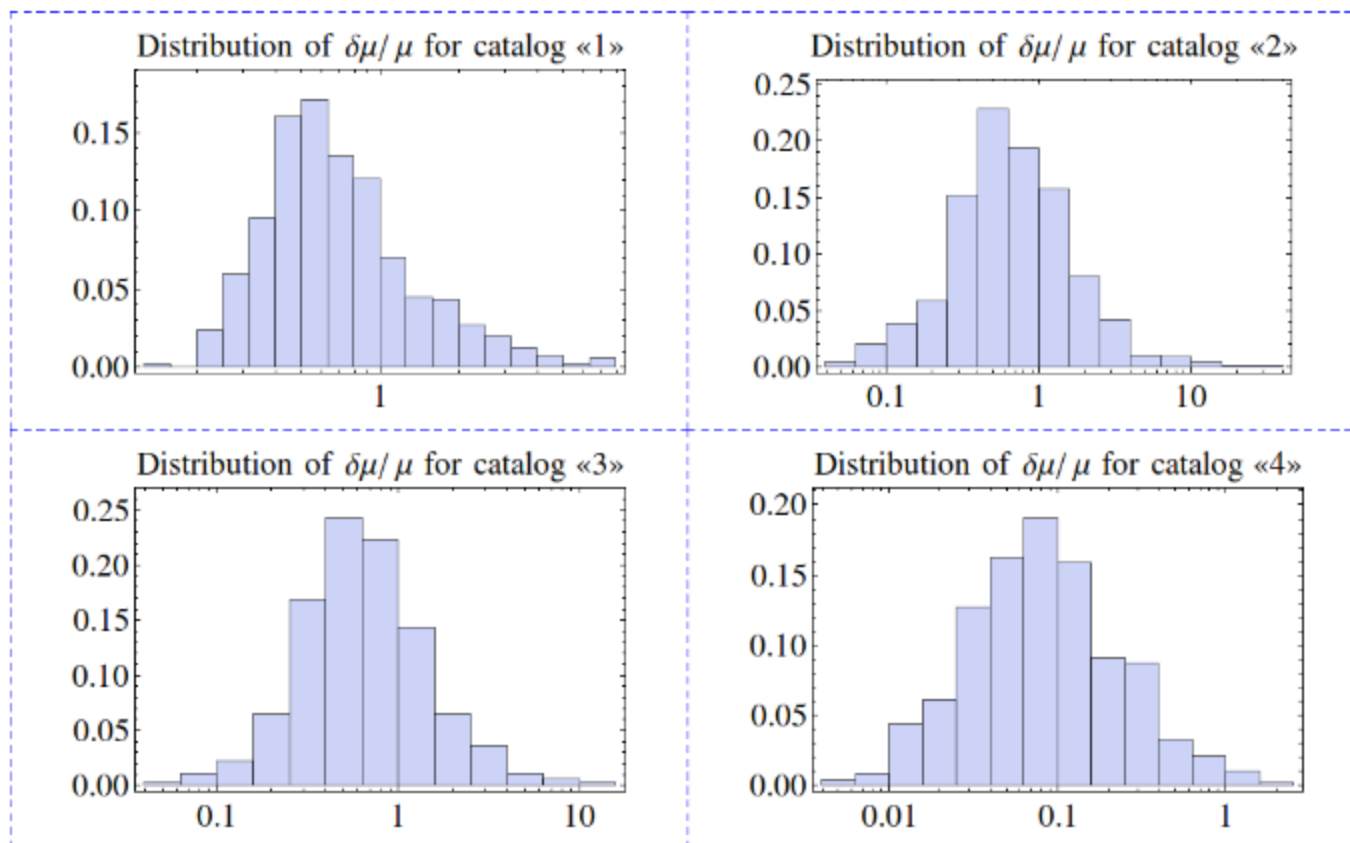


defining sources

candidates sources

other sources

Распределение ошибок скоростей



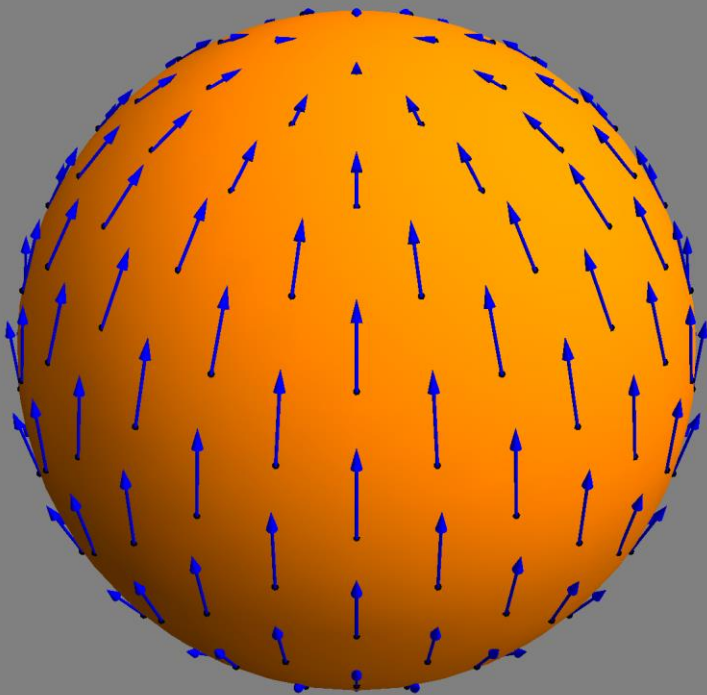
Representation of apparent motion

- **Decomposition of apparent velocity into a set of vector spherical harmonics**
- ***E* and *M* modes**

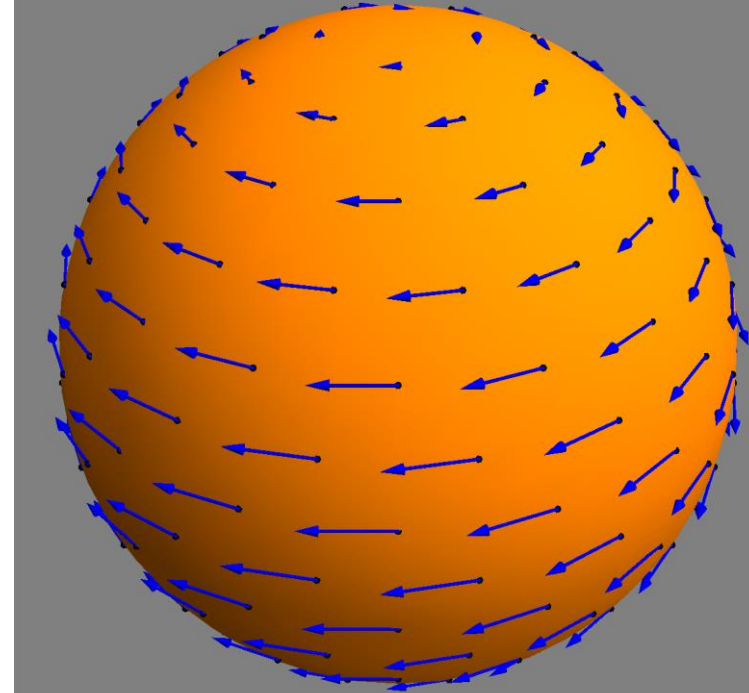
$$\vec{\mu}(\alpha, \delta) = \sum_{l,m} (a_{lm}^E \vec{Y}_{lm}^E + a_{lm}^M \vec{Y}_{lm}^M)$$

Dipole harmonic

$$\text{Re}(\bar{Y}_{n=1,m=0}^{l=1})$$



$$\text{Im}(\bar{Y}_{n=1,m=0}^{l=1})$$



Origin of apparent motion

Measurement errors

**Blandford – Rees effect
(superluminal motion)**

Accelerated motion of the Solar system

**Nonstationary and achromatic space-time
perturbations**

Accelerated motion of the Solar system

$$\vec{k} = \vec{K} + \frac{1}{c} \vec{K} \times [\vec{V} \times \vec{K}] + \frac{1}{c} \vec{K} \times [\vec{A} \times \vec{K}](t - t_0)$$

$$\vec{\mu} = \frac{1}{c} \vec{A} = \frac{1}{c} \frac{V^2}{R} \vec{n}$$

$\mu = 4 \mu as / yr$ - estimated

by Kopeikin and Makarov based on
modern model of our Galaxy

\vec{K} and \vec{k} are proper and apparent positions of a source

$$\vec{K} = (\sin \theta \cos \varphi, \sin \theta \sin \varphi, \cos \theta)$$

$$\vec{A} = (a_1, a_2, a_3)$$

$$\vec{\mu} = \frac{\partial \vec{k}}{\partial t} = \frac{1}{c} \vec{K} \times [\vec{A} \times \vec{K}]$$

in spherical basis $\{\vec{e}_r, \vec{e}_\theta, \vec{e}_\varphi\}$

$$\begin{aligned} \vec{\mu}(\theta, \varphi) = & \vec{e}_\theta (a_1 \cos \theta \cos \varphi + a_2 \cos \theta \sin \varphi - a_3 \sin \theta) \\ & + \vec{e}_\varphi (a_2 \cos \varphi - a_1 \sin \varphi) \end{aligned}$$

Apparent motion μ represent dipole spherical harmonic

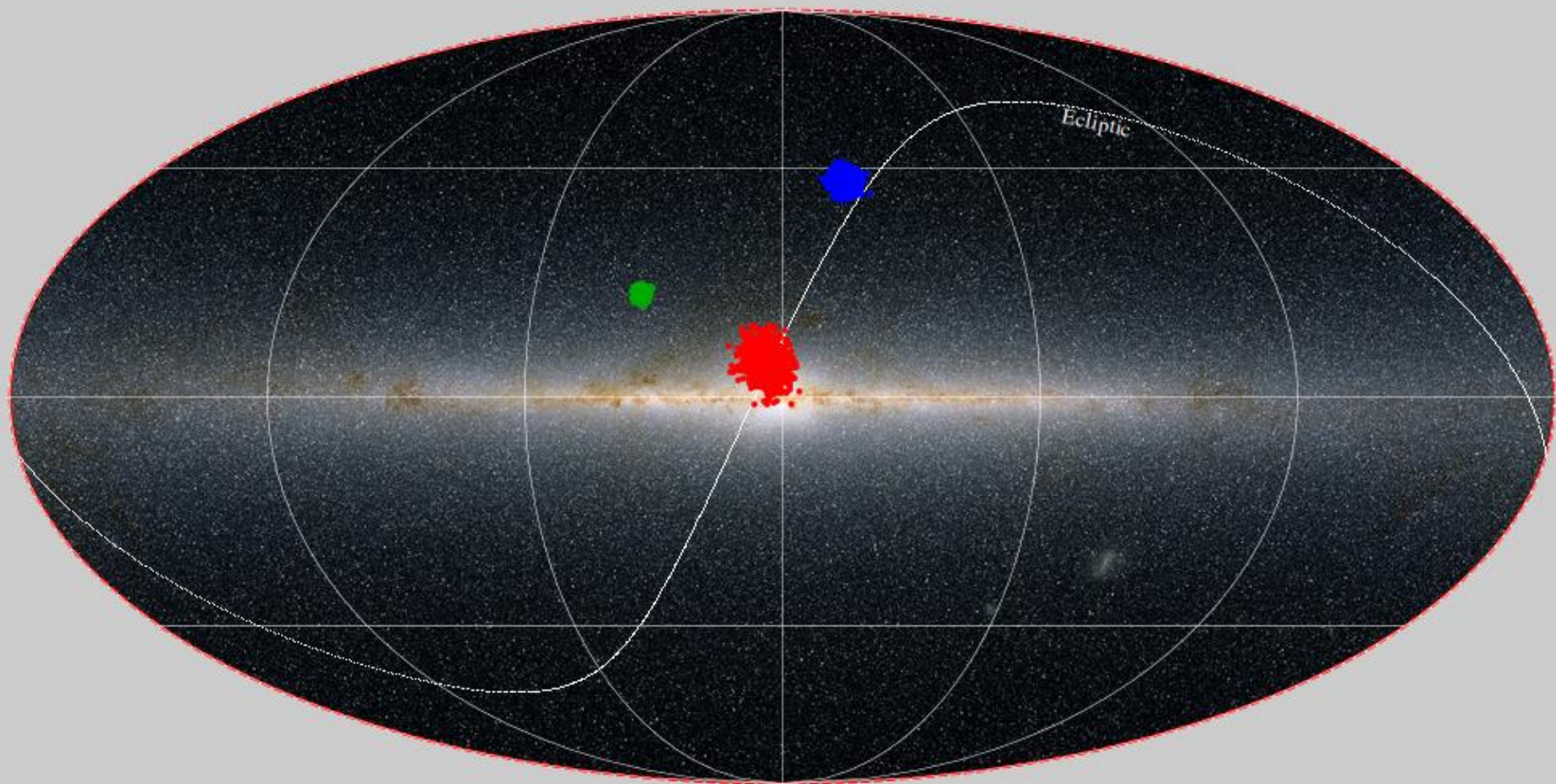
$$\vec{\mu} = \sum_{l=1}^{\infty} \sum_{m=-l}^{m=l} (a_{lm}^E \vec{Y}_{lm}^E + a_{lm}^M \vec{Y}_{lm}^M)$$

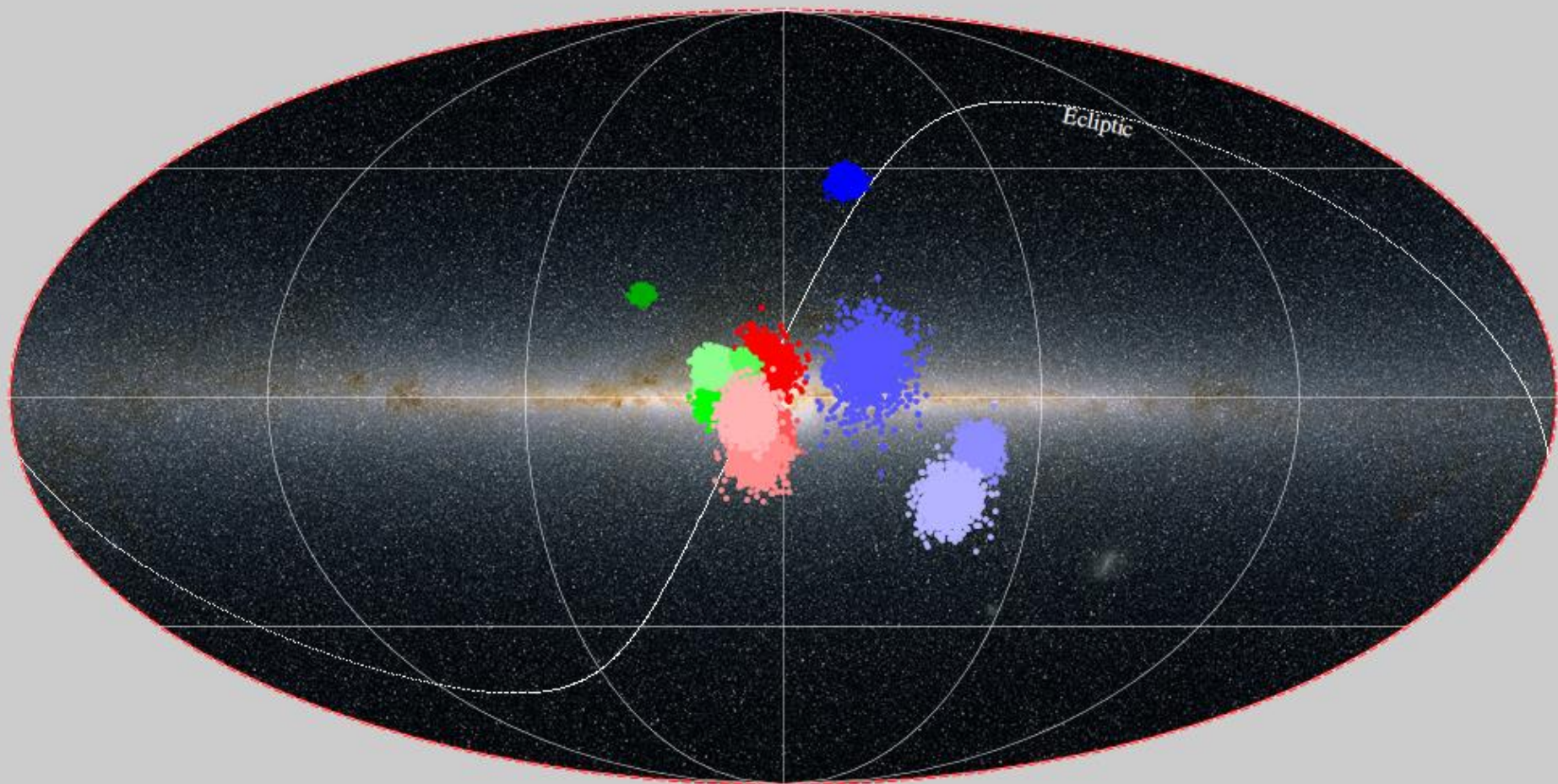
$$a_1 = \text{Im}(a_{1,1}^E) \sqrt{\frac{3}{4\pi}}$$

$$a_2 = \text{Re}(a_{1,1}^E) \sqrt{\frac{3}{4\pi}}$$

$$a_3 = \text{Re}(a_{1,0}^E) \sqrt{\frac{3}{8\pi}}$$

	cat 1	cat 2	cat 3
l=1	6.5 ± 0.3	16.3 ± 0.2	7.5 ± 0.2
l=1..2	7.2 ± 0.3	12.5 ± 0.3	3.6 ± 0.3
l=1..3	6.9 ± 0.3	13.8 ± 0.3	9.8 ± 0.3
l=1..4	7.3 ± 0.4	13.3 ± 0.3	$7. \pm 0.3$
	$\mu\text{as/year}$		





Value of secular apparent velocity measured is

$$\mu = 7 \pm 2 \mu as / yr$$

that coincides with result of Titov O. and
it is twice larger then expected
from theoretical assumption

Possible explanation:

**Noncircular motion of Solar system
inside our Galaxy**

**Dark object in vicinity of the Solar
system**

**MOND gravity for these scales (almost
perfect match)**

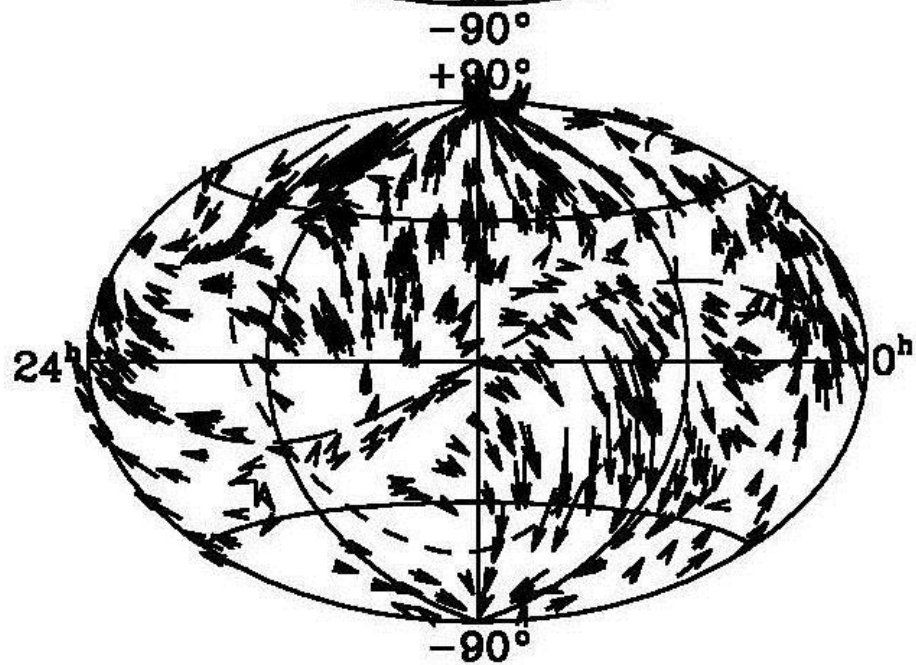
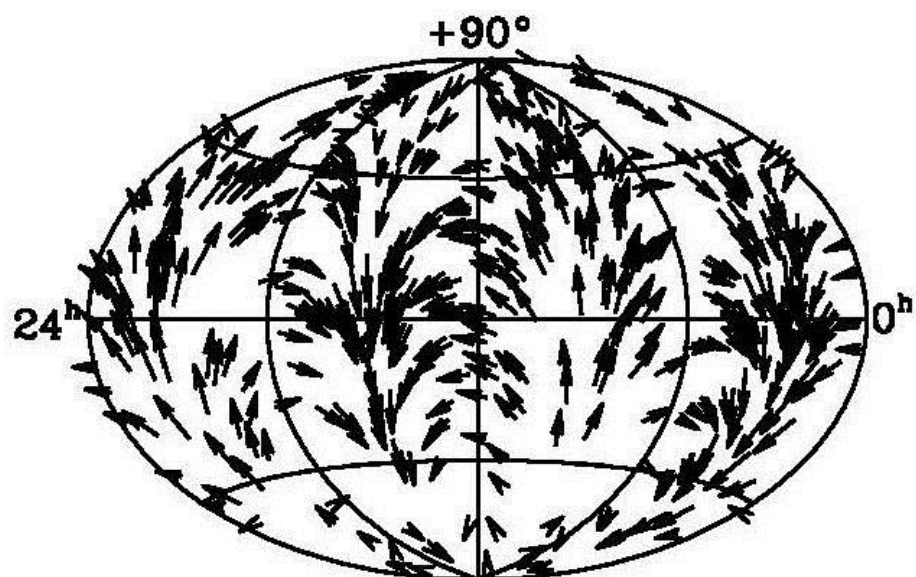
**Nonstationary and achromatic
space-time**

**Cosmological gravitational wave
background**

**Scalar perturbation induced by Dark
Energy fluctuations**

A gravitational wave traveling toward +z, with the “+” polarization, produces metric perturbations $h \cos \Omega t (\mathbf{x} \mathbf{x} - \mathbf{y} \mathbf{y})$ in the background coordinate reference frame, where h is the dimensionless strain of the wave, Ω is its angular frequency, and t is time. In the observer’s frame, the observed proper motion $\vec{\mu}$ of a radio source at position (θ, ϕ) will be

$$\vec{\mu} = \frac{1}{2} \Omega h \sin \Omega \eta \sin \theta (\vec{e}_\theta \cos 2\phi - \vec{e}_\phi \sin 2\phi)$$



$$\vec{\mu} = \Omega \, h \sin \Omega \eta \left[\frac{\sqrt{5}\pi}{6} (\vec{Y}_{22}^E + \vec{Y}_{2-2}^E - \vec{Y}_{22}^M + \vec{Y}_{2-2}^M) - \right. \\ \left. - \frac{\sqrt{70}\pi}{60} (\vec{Y}_{32}^E + \vec{Y}_{3-2}^E - \vec{Y}_{32}^M + \vec{Y}_{3-2}^M) \right] + \dots$$

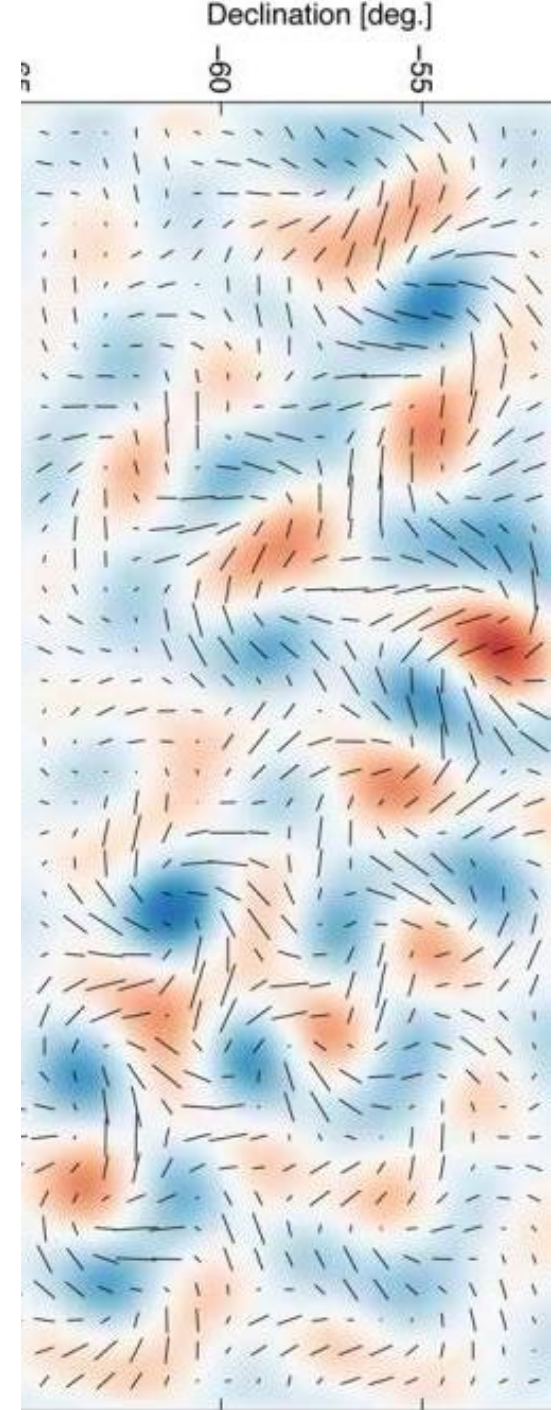
Cosmological Gravitational Waves

$k^{3/2}h(k) = A$ is an amplitude of a perturbation

$A_s \approx 5 \bullet 10^{-5}$ is the amplitude of scalar perturbation according WMAP measurements

$A_G = 0.2 \bullet A_s$ according to BICEP2 experiment

$A_G = \frac{1}{\sqrt{2\pi}} \frac{H_i}{m_{pl}}$ is predicted by theory(RSV, 1982)



Apparent angular velocity induced by Cosmological Gravitational Waves

$$\mu = \sqrt{\langle |\mu|^2 \rangle} = \sqrt{\int d^3k \frac{1}{4} c^2 k^2 \frac{A_G^2}{k^3}} \quad \text{is average angular velocity}$$

$$P_{\min} = 3 \bullet 10^8 \text{ yr}$$

$$\mu = 0.1 \frac{\mu\text{as}}{\text{yr}}$$

$$\text{Current accuracy is } 2 \frac{\mu\text{as}}{\text{yr}}$$



Thank you for attention

