

Black holes production in transplanckian collisions in (anti)-de Sitter spacetime

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(based on joint works with
I. Aref'eva, E. Guseva, and L. Joukovskaya,
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- In the case of two head-on colliding shock waves we derive and study the equation on formed trapped surface
- The volume of the trapped surface is being interpreted as the cross section of black holes production

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- Beyond the *AdS/CFT*-duality there is also *dS/CFT*-duality (Strominger - 2001). Therefore there is a chance to extend the holographic approaches to QCD on the de Sitter case
- *dS* background can be considered as the simplest model of spacetime with dark energy. In this context we analyze possible influence of dark energy on the process of black holes production in collisions

Metric of spacetime

- It is convenient to immerse D -dimensional $(A)dS$ spacetime with shock wave as submanifold to $(D + 1)$ -dimensional Minkowski space. In these terms the metric is following:

$$ds^2 = -dUdV + d\vec{Z}^2 + dZ^{D^2} + F(Z^D)\delta(U)dU^2,$$

$$\vec{Z} = \{Z^2, \dots, Z^{D-1}\}, \quad -UV + \vec{Z}^2 \pm Z^{D^2} = \pm a^2,$$

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- The shape of shock-wave could be obtained by boosting of the $(A)dS$ -Schwarzschild solution (Hotta, Tanaka - 1992):

$$\lim_{v \rightarrow 1} \frac{1}{\sqrt{1-v^2}} f\left(\frac{(Z^0 + vZ^1)^2}{1-v^2}; \lambda\right) = \delta(Z^0 + Z^1) \int_{-\infty}^{\infty} f(x^2; \lambda) dx$$

Shape function properties

- Explicit integral representation:

$$F(Z) = 2Ma^2 \cdot p.v. \int_{-\infty}^{\infty} \frac{(a^2(\pm Z^2 + x^2) + Z^2(x^2 \mp Z^2))}{(Z^2 \mp x^2)^2 (\pm a^2 + x^2 \mp Z^2)^{\frac{D-1}{2}}} dx$$

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- For example:

$$F_{4,dS}(Z^4) = 4MG_4 \left(-2 + \frac{Z^4}{a} \ln \left(\frac{1 + \frac{Z^4}{a}}{1 - \frac{Z^4}{a}} \right) \right)$$

$$F_{5,AdS}(Z^5) = \frac{3\pi MG_5}{2a} \left(\frac{2\frac{Z^5}{a^2} - 1}{\sqrt{\frac{Z^5}{a^2} - 1}} - 2\frac{Z^5}{a} \right)$$

Shape function properties

- In the charged case:

$$F_D(Q, Z) = F_D(Z) - \frac{Q^2}{2M} F_{2D-3}(Z),$$

where rescaled energy and charge are

$$M = \frac{8\pi G_D m}{(D-2)\Omega_{D-2}}, \quad Q^2 = \frac{8\pi G_D q^2}{(D-2)(D-3)}$$

Definition of trapped surface

- A trapped surface is a $(D - 2)$ -dimensional spacelike surface whose each two null normals have zero convergence (Neighboring light rays, normal to the surface, must move towards one another)

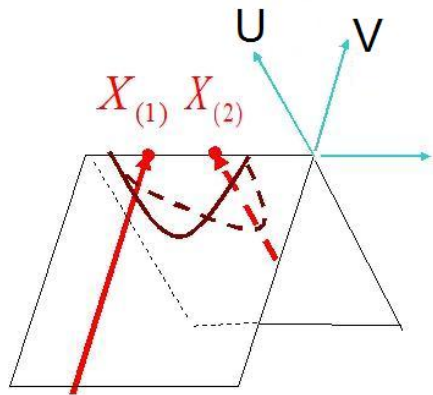
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- This has been proved for asymptotically flat spacetimes, but there is a common opinion that it is valid for dS/AdS too

Geometrical structure



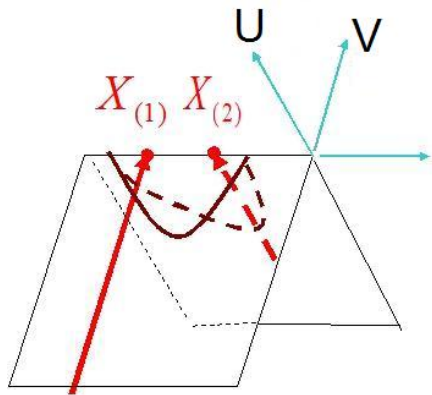
- Parts of trapped surface:

$$S = S_1 \cup S_2;$$

$$S_1 : U = 0, V = -\Psi_1(\rho),$$

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$$S_1 : U = 0, V = -\Psi_1(\rho),$$

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- On the surface of gluing:

$$\Psi_1 = \Psi_2$$

$$\left(\frac{d\Psi}{d\rho}\right)^2 = 4$$

Basic mathematical objects

- Trapped surface is a surface of null convergence:

$$\theta = h^{MN} \nabla_N \xi_M; \quad \theta = 0$$

Here:

ξ_M is a vector tangent to geodesic flow,

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- And (here $M, N = 0..D - 1$; $\alpha, \beta = 0..D - 3$):

$$h^{MN} = K_\alpha^M g^{\alpha\beta} K_\beta^N,$$

$$K_\alpha^M = (0, -\partial_\alpha \Psi, \delta_\alpha^M),$$

$$g_{\alpha\beta} = \frac{\partial Z^M}{\partial s^\alpha} \frac{\partial Z^N}{\partial s^\beta} g_{MN}$$

Dirichlet problem

- Equation on the trapped surface:

$$\left(\Delta_{\mathbb{S}^{D-2}/\mathbb{H}^{D-2}} \pm \frac{D-2}{a^2} \right) \frac{\psi - H}{1 \pm \frac{\rho^2}{2a^2}} = 0, \quad H = \frac{1}{2} \left(1 \pm \frac{\rho^2}{2a^2} \right) F$$

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- Boundary conditions in arbitrary dimension give us the following equation on the radius of the trapped surface:

$$\frac{1}{4} \left(1 \pm \frac{\rho_0^2}{2a^2} \right) F'(\rho_0) \pm \frac{\rho_0}{2a^2 \mp \rho_0^2} F(\rho_0) + 2 = 0$$

Criticality in de Sitter space

- In de Sitter space we deal with equation

$$\frac{1}{4}\left(1 + \frac{\rho_0^2}{2a^2}\right)F'(\rho_0) + \frac{\rho_0}{2a^2 - \rho_0^2}F(\rho_0) + 2 = 0$$

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- There is the explicit general formula:

$$\frac{(2 + x^2)^{D-2}}{x^{D-3}(2 - x^2)} = C_D \frac{a^{D-3}}{\rho}; \quad x = \rho/a$$

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- These functions have positive minima at the points

$$x_{min,D} = \frac{\sqrt{2(-1 + D - 2\sqrt{D-2})}}{\sqrt{D-3}}, \quad D > 3;$$

$$x_{min,3} = 0, \quad D = 3$$

When does the trapped surface exist in dS?

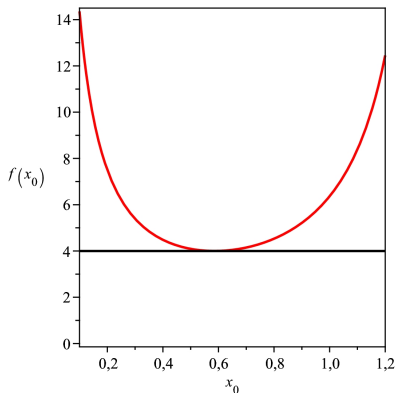


Figure: Critical effect of trapped surface non-formation.

Charged shock waves in dS

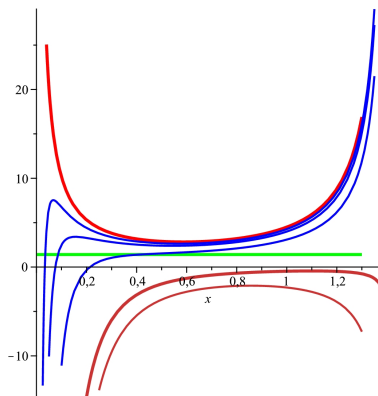


Figure: Influence of charge on the cross section of BH production in dS

Charged shock waves in AdS

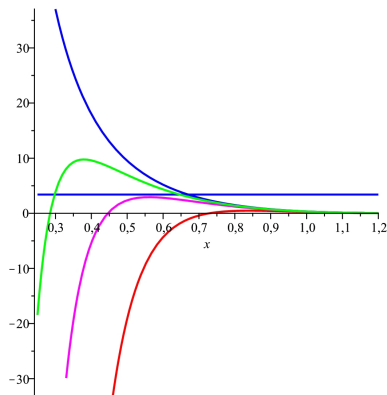


Figure: Influence of charge on the cross section of BH production in AdS

Volume of the trapped surface

- Volume of the trapped surface can be evaluated in general way and does not depend explicitly on the form of Ψ . We consider particular cases of dS_4 and AdS_5 both in the limit of low energy $M \ll a^{D-3}$:

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- In particular:

$$\mathcal{A}_{dS_4} \approx 4\pi\rho_0^2 \left(1 - \frac{\rho_0^2}{2a^2}\right) \approx 32\pi M^2 \left(1 - \frac{4M^2}{a^2} - \frac{3\pi}{64} \frac{Q^2}{M^2}\right)$$

$$\mathcal{A}_{AdS_5} \approx 4\pi a^3 \left(\frac{3\pi}{2} \frac{M}{a^2}\right)^{2/3} \left(1 - \frac{1}{24} \left(1 + \frac{5Q^2}{Ma^2}\right) \left(\frac{4\sqrt{2}}{3\pi} \frac{a^2}{M}\right)^{2/3}\right)$$

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- For dS spacetime it was shown that if energy density of neutral shock waves is very large or if the spacetime is strongly curved then black hole could not be formed even in head-on collision. For $D = 3$ the critical value of energy density does not depend on the cosmological radius
- Influence of charge on the cross section of black holes production has been calculated

Outlook

- Analysis of shock waves with more complicated inner structure (spin, scalar phantom charge etc.)

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- Holographic description. Application of shock wave technics to solution of QCD problems (Gubser, Yarom, Pufu - 2008)