The Arrow of Time and Effective Gravity in Modified Spinor Quantum Dynamics

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1. Universe as Ensemble of Fermionic Universes

Multi Universe wave function

$$\Psi = \begin{pmatrix} \Psi_1 \\ \Psi_2 \\ \Psi_3 \\ \vdots \end{pmatrix}$$

Extended Dirac Equation with the complex parameter of evolution:

$$i\hbar\Psi_{,\tau} = \widehat{H}\Psi = (\widehat{\varepsilon} - \frac{i\hbar}{2}\widehat{\Im})\Psi, \qquad \Psi_{,\tau^*} = 0$$

 $\tau = t - \frac{i\hbar}{2}\beta$

$$\widehat{\varepsilon} = \begin{pmatrix} E & 0 & 0 & & \\ 0 & E & 0 & \cdots & \cdot \\ 0 & 0 & E & & \\ & \ddots & & \ddots & \cdot \end{pmatrix}, \quad \widehat{\Im} = \begin{pmatrix} \gamma_1 & 0 & 0 & & \\ 0 & \gamma_2 & 0 & \cdots & \cdot \\ 0 & 0 & \gamma_3 & & \\ & \ddots & & \ddots & \cdot \end{pmatrix}$$
$$\gamma_1 < \gamma_2 < \gamma_3 < \cdots = const$$

Matrix element for arbitrary operator \hat{A} :

$$<\Psi_1|\hat{A}|\Psi_2>=\int \overline{\Psi}_1 \gamma^4 \hat{A} \Psi_2 d^4 x$$

Average value

$$\bar{\hat{A}} = \frac{1}{Z} \int \bar{\Psi} \gamma^4 \hat{A} \Psi d^4 x$$

Statistical sum

$$Z(\beta) = \langle \Psi | \Psi \rangle = \int \overline{\Psi} \gamma^4 \Psi \, d^4 x - not \, a \, constant$$

Global Arrow of Time – only one Universe will survive asymptotically

$$\frac{d\overline{\mathfrak{I}}}{dt} \le 0$$

2. Dynamic Reduction of the Additional Coordinates

Fully Analytic Model with Extra Dynamic Parameters

$$X^M = x^M - \frac{\mathrm{i}\hbar}{2}\beta^M$$

Master Dirac Equation

$$\Psi_{(k),X^4} + \left((\gamma^4)^{-1} \gamma^\mu \frac{\partial}{\partial X^\mu} + \frac{1}{2} \gamma_k \right) \Psi_{(k)} = 0, \qquad \Psi_{(k),X^{M*}} = 0$$

$$\Psi = \Psi(X^{M}) = \begin{pmatrix} \Psi_{(1)} \\ \vdots \\ \vdots \end{pmatrix} + \begin{pmatrix} \Psi_{(2)} \\ \vdots \\ \vdots \end{pmatrix} + \begin{pmatrix} \Psi_{(3)} \\ \vdots \\ \psi_{(n)} \end{pmatrix} + \dots + \begin{pmatrix} \vdots \\ \Psi_{(n)} \end{pmatrix}$$

Extreme Principle and Probabilistic Reduction of Dynamic Parameters

$$\rho = \rho(x^{N}, \beta^{M})$$
$$\max_{\beta^{M}x^{4}} \rho = \rho(x^{\mu}, x^{4}_{max}(x^{\mu}), \beta^{M}_{max}(x^{\mu}))$$
$$\beta^{M} = \beta^{M}_{max}(x^{\mu}) \quad \& \quad x^{4} = x^{4}_{max}(x^{\mu})$$

Thermodynamic mode = The evolution of the ensemble of Universes

Dynamically Reduced Wave Function

$$\psi(x^{\mu}) = \Psi[X^L(x^{\mu})]$$

Reduced Master Dirac Equation (for each Universe)

$$\Psi_{,x}^{\mu} = (D^{-1})^{\mu}_{\nu} \psi_{,x}^{\mu}$$
$$D^{\mu}_{\nu} = X^{\mu}_{,x^{\nu}} - X^{4}_{,x^{\nu}} (\gamma^{4})^{-1} \gamma^{\mu}$$

3. Effective Clifford Gravity

$$m_{tot}^2 = p^2 \left(X^{\lambda}(x^{\mu}) \right) = \frac{1}{J^4} \overline{\Psi} \gamma^4 \hat{p}^2 \Psi$$
$$m_{geom}^2 = p^2 (x^{\mu}) = \frac{1}{j^4} \overline{\Psi} \gamma^4 \hat{p}_{\mu} \hat{g}^{\mu\nu} \hat{p}_{\nu} \Psi$$
$$m_{NR}^2 = m_{tot}^2 - m_{geom}^2$$

$$\hat{g}^{\mu\nu} = \frac{1}{2} \eta^{\alpha\beta} \left\{ [(D^{-1})^{\mu}_{\alpha}]^{+} \gamma (D^{-1})^{\nu}_{\beta} + [(D^{-1})^{\nu}_{\alpha}]^{+} \gamma (D^{-1})^{\mu}_{\beta} \right\}$$
$$\gamma = \gamma^{0} \gamma^{4}$$
$$g^{\mu\nu}(x^{\mu}) = \frac{1}{j^{4}} \overline{\psi} \gamma^{4} \widehat{g}^{\mu\nu} \psi$$

Metrics operator combines impact of all Universes, but asymptotically metric becomes metric of one survived Universe.

Rules for Riemannian tensors operations:

$$\hat{g}^{\mu\nu}\hat{g}_{\nu\lambda} = \delta^{\mu}_{\nu}\hat{I}$$
$$\hat{g}_{\mu\lambda} = a_{\mu\nu}I + b_{\mu\nu}\gamma^{5} + c_{\mu\nu|\alpha}\gamma^{\alpha} + d_{\mu\nu|\alpha}\lambda^{\alpha} + e_{\mu\nu|\alpha\beta}\Sigma^{\alpha\beta}$$

$$\nabla_{\mu}\hat{A}^{\nu} = \partial_{\mu}\hat{A}^{\nu} + \left\{\widehat{\Gamma}^{\nu}_{\mu\lambda}, \hat{A}^{\lambda}\right\}$$
$$\widehat{\Gamma}^{\tau}_{\lambda\nu} = \left\{\widehat{\Gamma}_{\lambda\nu|\sigma}, \hat{g}^{\sigma\tau}\right\} = \left\{\partial_{\lambda}(\hat{g}_{\sigma\nu}) + \partial_{\nu}(\hat{g}_{\lambda\sigma}) - \partial_{\sigma}(\hat{g}_{\nu\lambda}), \hat{g}^{\sigma\tau}\right\}$$

$$\begin{split} \widehat{R}^{\lambda}_{\mu\nu\rho} &= \partial_{\mu}\widehat{\Gamma}^{\lambda}_{\nu\rho} - \partial_{\nu}\widehat{\Gamma}^{\lambda}_{\mu\rho} + \left\{\widehat{\Gamma}^{\lambda}_{\mu\tau}, \widehat{\Gamma}^{\tau}_{\nu\rho}\right\} - \left\{\widehat{\Gamma}^{\lambda}_{\nu\tau}, \widehat{\Gamma}^{\tau}_{\mu\rho}\right\} \\ R^{\lambda}_{\mu\nu\rho}(x^{\mu}) &= \frac{1}{j^{4}}\overline{\psi}\gamma^{4}\widehat{R}^{\lambda}_{\mu\nu\rho}\psi \end{split}$$

4. Cosmological Solution

$$\Psi_{,x^i} = 0, \qquad i = 1,2,3$$

General solution:

$$\Psi = \Psi_{+} + \Psi_{-}, \qquad \Psi_{\pm} = \frac{1}{\sqrt{2}} \begin{pmatrix} f_{\pm} \\ \mp i f_{\pm} \end{pmatrix}$$

and 2-columns $f_{+} = f_{+} (X^{0} \mp X^{4})$

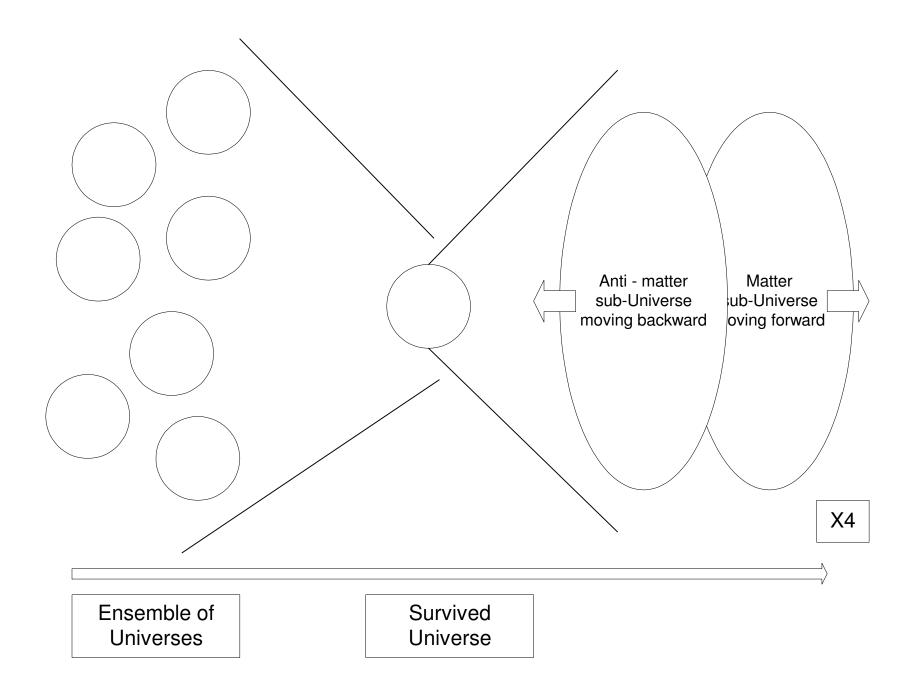
Thermodynamic modes are:

$$\max_{\beta^4,\beta^0,x^4} \rho = \rho(x^0, x^4_{max}(x^0), \beta^0_{max}(x^0), \beta^4_{max}(x^0))$$
$$x^4 = x^4_{max}(x^0) \quad \& \quad \beta^4 = \beta^4_{max}(x^0) \quad \& \quad \beta^0 = \beta^0_{max}(x^0)$$

Particularly asymptotically for $x^4 \rightarrow \infty$:

$$x^4 \approx x^0$$

Hamiltonian time becomes equal to Minkowskian time of the survived Universe



Metrics

$$\hat{g}^{\mu\nu} = \left[\frac{1 - (\dot{x}^4)^2 + 2(\dot{x}^4)^4 - 2(\dot{x}^4)^3\hat{\Gamma}^0}{(1 - (\dot{x}^4)^2)^2}\eta^{\mu\nu} + \frac{2}{1 - (\dot{x}^4)^2}(\dot{x}^4)^2\delta_0^{\mu}\delta_0^{\nu}\right] + \frac{1 + (\dot{x}^4)^2}{(1 - (\dot{x}^4)^2)^2}\dot{x}^4(\delta_0^{\mu}\hat{\Gamma}^{\nu} + \delta_0^{\nu}\hat{\Gamma}^{\mu})$$

For

 $\dot{x}^4\approx 1$

$$\hat{g}_{\mu\nu} = \frac{1}{6} \left[\frac{3}{2} (\delta^{0}_{\mu} \hat{\Gamma}_{\nu} + \delta^{0}_{\nu} \hat{\Gamma}_{\mu}) \dot{x}^{4} - 2(\dot{x}^{4})^{2} \hat{\Gamma}_{\mu} \hat{\Gamma}_{\nu} - \eta_{\mu\nu} \right] (1 + \dot{x}^{4} \hat{\Gamma}^{0})$$
$$g_{\mu\nu}(x^{0}) = \frac{1}{j^{4}} \overline{\psi} \gamma^{4} \hat{g}_{\mu\nu} \psi$$

$$ds^2 = g_{\mu\nu}(x^0)dx^{\mu}dx^{\nu} = 2\omega_k dx^0 dx^k + \Xi \eta_{kl} dx^k dx^l$$

5. Conclusions

- a. Multi Universe ensemble approach proposed as description of the Nature.
- b. For such model the Arrow of Time appears as evolution from mixed initial data with all Universes to one Universe with the longest life time.
- c. Five dimensional Dirac's equation with complex parameters plays a role of master equation for this system.
- d. Dynamic reduction mechanism based maximum of the probability density defines evolution of ensemble of Universes. Arrow of Time defines which of them will survive.
- e. Reduced theory is strongly non-linear with fixed parameters of anisotropy.
- f. Metrics can be introduced through "geometrization" of mass distribution. Metrics is an operator based on Clifford algebra. All non-operator Riemannian structures can be calculated as coordinate realizations of the Riemannian ones.
- g. Cosmological case is considered with Universe which consists of matter sub-Universe moving forward in time and anti-matter sub-Universe moving backward in time.